

# Rational Sunspots

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## **Nonlinear Models in Macroeconomics and Finance for an Unstable World**

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<sup>1</sup>The views expressed are solely the responsibility of the authors and should not to be interpreted as reflecting the views of Sveriges Riksbank.

# This Paper

- 1 Propose a method to consider a broader class of solutions to stochastic linear models. Two generalizations:
  - A novel way to introduce sunspots that yields drifting parameters and stochastic volatility
  - Include temporary unstable solutions: we allow for determinacy, indeterminacy and instability
- 2 Develop an econometric strategy to verify if unstable paths are empirically relevant
- 3 Application:
  - Example of U.S. Great inflation (Lubik and Schorfheide, 2004, model and data)
  - U.S. inflation dynamics in the 70's are better described by unstable equilibrium paths.

# Motivation

- RE generally implies multiple equilibria
  - Explosive
  - Stable
- How can we get uniqueness? (Sargent and Wallace ,1973; Phelps and Taylor, 1977; Taylor, 1977; Blanchard, 1979)
- *Stability Criterion*: Transversality conditions  
In saddle paths dynamics only one solution is stable
- This became the standard in Macroeconomics (Blanchard and Kahn, 1980)

## Example: U.S. Great Inflation period

*Is it appropriate to rule out unstable paths from the empirical analysis?*

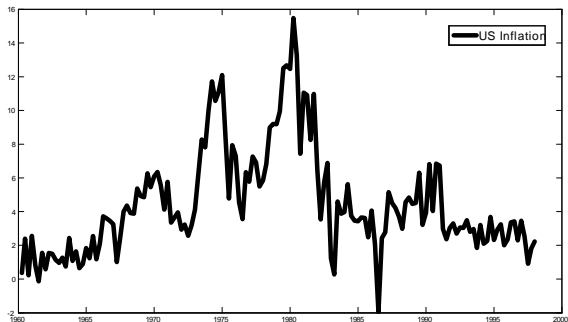


Figure: CPI inflation, quarterly data. Sample: 1960Q1 - 1997Q4

*Is there any evidence that inflation is described (at least for a while) by unstable equilibria?*

## A simple example: multiple RE solutions

Consider the following simple one equation model:

$$y_t = \frac{1}{\theta} E_t y_{t+1} + \varepsilon_t, \quad \varepsilon_t \sim N(0, \sigma_\varepsilon^2) \quad (1)$$

Equation (1) has an infinite number of solutions:

$$\begin{aligned} y_{t+1} &= E_t y_{t+1} + \eta_{t+1} \\ y_{t+1} &= \theta y_t - \theta \varepsilon_t + \eta_{t+1} \end{aligned} \quad (2)$$

where  $E_t \eta_{t+1} = 0$ .

Assume:

$$\eta_{t+1} = (1 + M)\varepsilon_{t+1} + \zeta_{t+1}$$

where  $\zeta_{t+1}$  = sunspot or non-fundamental error.

Two sources of multiplicity:

*This paper considers the FIRST term: intrinsic multiplicity of RE solutions*

## A simple example: multiple RE solutions

Assume  $\zeta_t = 0 \forall t$ . All the solutions for  $y_t$  are described by

$$y_t = \theta y_{t-1} - \theta \varepsilon_{t-1} + (1 + M) \varepsilon_t \quad (3)$$

Degree of freedom: the solution is parameterized by  $M \in (-\infty, +\infty)$

- Two famous particular cases:

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- Two famous particular cases:
  - "pure" forward-looking solution:  $M = 0$  ( $\eta_t = \varepsilon_t$ )

$$\begin{aligned} y_t^F - \theta y_{t-1}^F &= \varepsilon_t - \theta \varepsilon_{t-1} \\ y_t^F &= \varepsilon_t \quad \forall t \end{aligned} \quad (4)$$

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- "pure" backward-looking solution ( $M = -1$ )

$$y_t^B = \theta y_{t-1}^B - \theta \varepsilon_{t-1} \quad (5)$$



# The interpretation for M

- For  $M \neq -1$ , the expected value = an exponentially weighted average of the past observations (Muth, 1961)

$$E_t y_{t+1} = M \sum_{i=0}^t \left( \frac{\theta}{1+M} \right)^i y_{t+1-i}$$

*Natural interpretation for M: the way agents form expectations*

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*Natural interpretation for  $M$ : the way agents form expectations*

- $M$  defines the importance the agents give to the past data, both in *absolute* terms ( $M$  vs 0), and in *relative* terms.

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## *Natural interpretation for M: the way agents form expectations*

- $M$  defines the importance the agents give to the past data, both in *absolute* terms ( $M$  vs 0), and in *relative* terms.
- Infinite solutions = infinite way we can set that weights  $\Rightarrow$  how to choose?

# The stability criterion (e.g., Blanchard and Kahn, 1980)

$$y_t = \theta y_{t-1} - \theta \varepsilon_{t-1} + (1 + M) \varepsilon_t$$

Is the stability criterion sufficient to identify a unique path?

① If  $|\theta| > 1$     **YES**    determinacy, by imposing  $M = 0 =$  f.l.  
solution  $y_t^F = \varepsilon_t$  (MSV solution)

② If  $|\theta| < 1$     **NO**    indeterminacy

$\Rightarrow$  *"Sunspot equilibria can often be constructed by randomizing over multiple equilibria of a general equilibrium model"* Benhabib and Farmer (1999, p.390)

# Introducing sunspot equilibria

We have infinite equilibria because:

- there is an infinite number of ways of forming expectations
- parametrized by  $M$

hence we introduce sunspots randomizing over  $M$ :

$$M_t = M_t(\zeta_t) \quad (6)$$

where  $\zeta_t$  i.i.d., orthogonal to the fundamental shocks  $\varepsilon_s$  ( $s = 1, 2, \dots$ ), and  $E_t \zeta_t = 0 \forall t$ .

## Introducing sunspot equilibria: drifting parameters and unstable paths

If  $M_t$  is a stochastic process with  $E_t M_{t+1} = M_t$  then

$$y_t = \alpha_t y_{t-1} - \alpha_t \varepsilon_{t-1} + (1 + M_t) \varepsilon_t \quad (7)$$

with  $\alpha_t = \theta \frac{M_t}{M_{t-1}}$  (with  $M_{t-1} \neq 0$  otherwise FL solution).

- Same form as  $y_t = \theta y_{t-1} - \theta \varepsilon_{t-1} + (1 + M) \varepsilon_t$

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- Drifting parameters and stochastic volatility within the rational expectations framework. Cogley and Sargent (2005), Primiceri (2005).
- Intuition: agents can modify in every period the expectation formation process
- *Reconsidering unstable paths:  $|\theta| > 1$  and  $M_t$  temporarily different from zero*

## Which process for M? RE Solutions

- With  $M_t$  random variable the forecast error becomes:

$$\eta_t = (1 + M_t)\varepsilon_t + (M_t - M_{t-1}) \left( \sum_{i=1}^{t-1} \theta^i \varepsilon_{t-i} \right) \quad (8)$$

- Under RE  $E_{t-1}\eta_t = 0$ , then:

- 1  $E_{t-1}(M_t) = M_{t-1}$  ( $M_t$  martingale)
- 2  $E_{t-1}[(1 + M_t)\varepsilon_t] = 0$ , ( $M_t$  must be uncorrelated with  $\varepsilon_t$ )

if  $|\theta| < 1$  : Use conditions 1 and 2

if  $|\theta| > 1$  : Unstable paths. Consider the role of transversality condition

## Temporarily explosive paths

- All the paths corresponding to expected temporary deviations of  $M_t$  from 0 will not violate the transversality condition
- RE requires  $M_t$  martingale, but if  $|\theta| > 1$  and  $M_t$  martingale, when  $M_t \neq 0$  the economy is expected to remain on the unstable path, so transversality condition would be violated

$\Rightarrow$  To allow for temporary unstable paths relax the martingale assumption (and RE)

Deviation can be minimal without practical implications

► Mprocess

## Example: Lubik and Schorfheide (2004) model

$$x_t = E_t(x_{t+1}) - \tau(R_t - E_t(\pi_{t+1})) + g_t$$

$$\pi_t = \beta E_t(\pi_{t+1}) + \kappa(x_t - z_t)$$

$$R_t = \rho_R R_{t-1} + (1 - \rho_R)(\psi_1 \pi_t + \psi_2(x_t - z_t)) + \varepsilon_{R,t}$$

and

$$g_t = \rho_g g_{t-1} + \varepsilon_{g,t}; \quad z_t = \rho_z z_{t-1} + \varepsilon_{z,t}$$

allow for non-zero correlation between the two shocks:  $\rho_{gz}$

Compare two "models":  $M_S$  (stable solutions) and  $M_U$  (unstable solutions).

► Eig

# The estimation strategy

- The model has stochastic volatility, then the likelihood distribution is not Gaussian
- Common practice in non linear DSGE literature: use particle filter to approximate the likelihood in MCMC (Fernandez-Villaverde and Rubio-Ramirez, 2007)
- **Different approach:** Particle filter to approximate the posterior distribution of the parameters and  $M_t$
- Sequential Learning on the parameters: how inference evolves over time gives additional information about the role of sunspots and unstable paths

# Particle filter

- 1 Marginalization.  $l$ : all latent states different from  $M$ ;  $y$ : data

$$p(l, M|y) = \underbrace{p(l|y, M)}_{\text{Kalman Filter}} \underbrace{p(M|y)}_{\text{Particle Filter}}$$

- 2 Parameter learning, combining:
  - 1 Particle Learning by Carvalho, Johannes, Lopes and Polson (2010):
  - 2 Liu and West (2001)

See also Chen, Petralia and Lopes (2010)

# Priors and Distributions: same as in LS

Table 1: Prior Distributions

Parameter	Density	Mean	Standard Deviation
$\psi_1$	Gamma	1.1	0.5
$\psi_2$	Gamma	0.25	0.15
$\rho_R$	Beta	0.5	0.2
$\pi^*$	Gamma	4	2
$r^*$	Gamma	2	1
$\kappa$	Gamma	0.5	0.2
$\tau^{-1}$	Gamma	2	0.5
$\rho_g$	Beta	0.7	0.1
$\rho_z$	Beta	0.7	0.1
$\gamma$	Beta	0.8	0.15
$\sigma_R$	Inverse Gamma	0.31	0.16
$\sigma_\zeta$	Inverse Gamma	0.1	0.05
Variance Covariance	Density	Scale	Degrees of freedom
$\Sigma_{gz}$	Inverse Wishart	3 $\begin{bmatrix} 0.38^2 & 0 \\ 0 & 1 \end{bmatrix}$	8

# Estimates Great Inflation sample

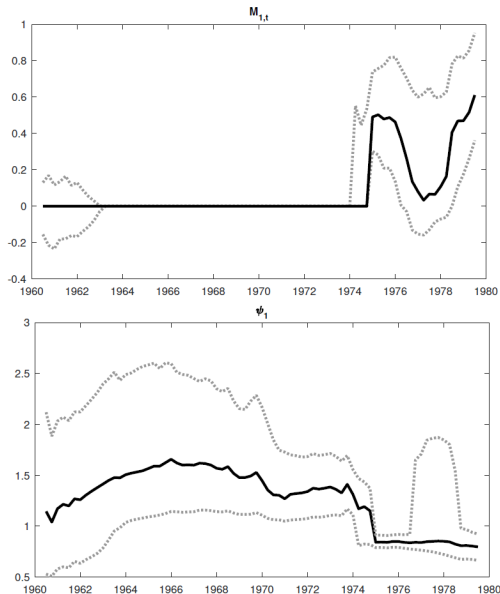
Parameter	Pre- Volcker 1960:I - 1979:II		
	$M_S$	$M_U$	$LS$
$\psi_1$	0.80 [0.66 0.92]	0.76 [0.61 0.91]	0.77 [0.64 0.91]
$\psi_2$	0.16 [0.11 0.20]	0.20 [0.16 0.34]	0.17 [0.04 0.30]
$\rho_R$	0.68 [0.65 0.71]	0.60 [0.53 0.68]	0.60 [0.42 0.78]
$\pi^*$	1.90 [1.62 2.25]	1.73 [1.31 2.47]	4.28 [2.21 6.21]
$r^*$	1.41 [1.29 1.58]	1.23 [0.93 1.74]	1.13 [0.63 1.62]
$\kappa$	0.14 [0.10 0.18]	0.10 [0.07 0.14]	0.77 [0.39 1.12]
$\tau^{-1}$	3.41 [2.65 4.51]	3.02 [2.46 3.74]	1.45 [0.85 2.05]
$\rho_g$	0.64 [0.59 0.69]	0.68 [0.63 0.74]	0.68 [0.54 0.81]
$\rho_z$	0.76 [0.72 0.80]	0.75 [0.67 0.81]	0.82 [0.72 0.92]
$\rho_{gz}$	0.26 [0.19 0.37]	0.16 [0.06 0.25]	0.14 [-0.4 0.71]
$\gamma$	—	0.96 [0.85 0.99]	—
$\sigma_R$	0.22 [0.2 0.26]	0.19 [0.16 0.22]	0.23 [0.19 0.27]
$\sigma_g$	0.35 [0.3 0.4]	0.31 [0.24 0.37]	0.27 [0.17 0.36]
$\sigma_z$	1.11 [0.97 1.29]	1.00 [0.85 1.31]	1.13 [0.95 1.30]
$\sigma_\zeta$	0.08 [0.07 0.1]	0.06 [0.05 0.08]	0.20 [0.12 0.27]

90% credibility interval in brackets



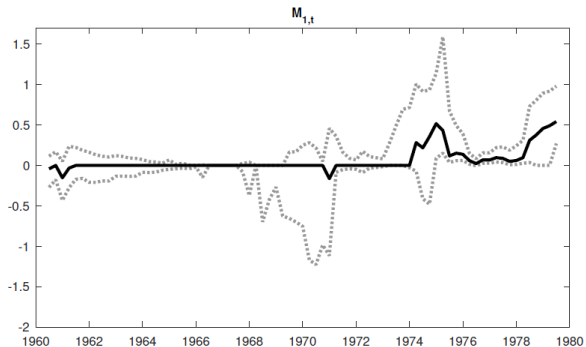
## Stable Model, Great Inflation:

Estimated path for  $M_t$  and sequential inference on the parameter  $\psi_1$ .



# Unstable Model, Great Inflation

The behavior of  $M_t$

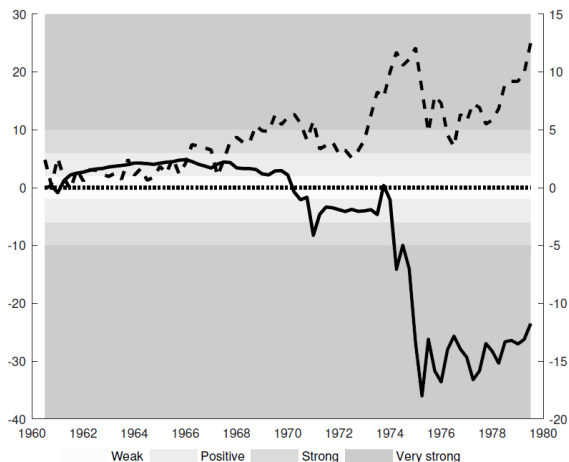


► IRFs

► Models

# Comparing the relative fit of $M_s/M_u$

Cumulative Bayes Factor:  $2 \ln(W_t)$  and the inflation rate



The Bayes Factor strongly favours the unstable model

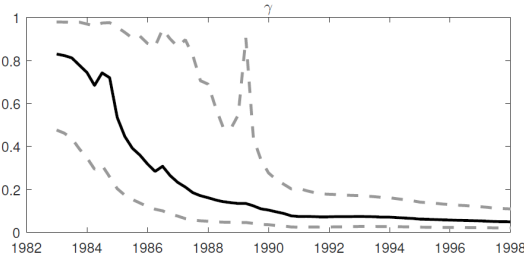
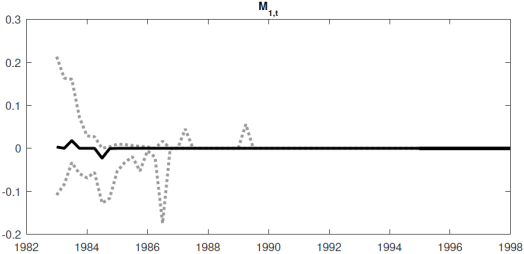
# Estimates Great Moderation sample

Parameter	Post-82 1982:IV - 1997:IV		
	$M_S$	$M_U$	$LS$
$\psi_1$	2.07 [1.50 2.81]	2.12 [1.07 3.65]	2.19 [1.38 2.99]
$\psi_2$	0.16 [0.06 0.39]	0.51 [0.19 1.38]	0.30 [0.07 0.51]
$\rho_R$	0.86 [0.81 0.90]	0.85 [0.80 0.89]	0.84 [0.79 0.89]
$\pi^*$	3.18 [2.73 3.74]	3.22 [2.74 3.74]	3.43 [2.84 3.99]
$r^*$	2.65 [1.87 3.56]	2.81 [2.17 3.60]	3.01 [2.21 3.80]
$\kappa$	0.28 [0.20 0.40]	0.64 [0.38 1.09]	0.58 [0.27 0.89]
$\tau^{-1}$	2.52 [1.90 3.41]	1.28 [0.85 1.88]	1.86 [1.04 2.64]
$\rho_g$	0.77 [0.68 0.84]	0.81 [0.74 0.87]	0.83 [0.77 0.89]
$\rho_z$	0.71 [0.62 0.80]	0.78 [0.70 0.84]	0.85 [0.77 0.93]
$\rho_{gz}$	0.03 [0.00 0.07]	0.04 [0.01 0.08]	0.36 [0.06 0.67]
$\gamma$	—	0.05 [0.02 0.11]	—
$\sigma_R$	0.16 [0.13 0.19]	0.16 [0.14 0.2]	0.18 [0.14 0.21]
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$\sigma_\zeta$	—	—	—

90% credibility interval in brackets

# Unstable Model, Great Inflation

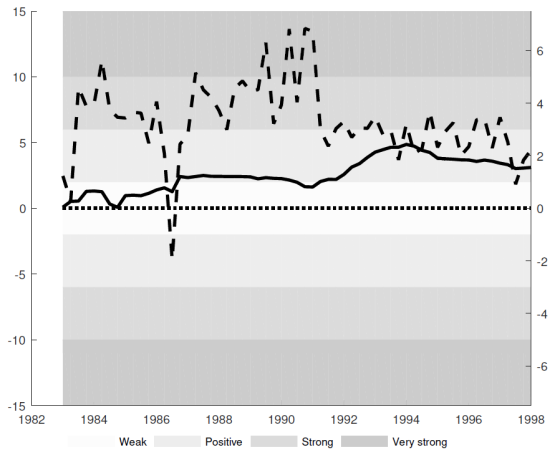
Estimated path for  $M_t$  and sequential inference on the parameter  $\gamma$ .



Model  $M_U$  selects the unique stable solution

# Comparing the relative fit of $M_s/\mu$

Cumulative Bayes Factor:  $2 \ln(W_t)$  and the inflation rate



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- Our methodology allows the data to choose between different possible alternatives: determinacy, indeterminacy and instability
- When the data are allowed this possibility, they unambiguously select the unstable model to explain the stagflation period in the '70s
- **Temporary unstable paths** can be empirically relevant and should not be excluded a priori

# EXTRA

## Which process for M? Temporary unstable paths

To guarantee the transversality condition relax the martingale assumption (and RE):

$$M_t = N_t A_{t-1}$$

- $N_t$  martingale
- $A_t \in \{0, 1\}$  non increasing random sequence

Indicate with  $\bar{T}$  random variable:  $\bar{T} = \inf \{t : A_t = 0\}$

Properties of  $M_t$  :

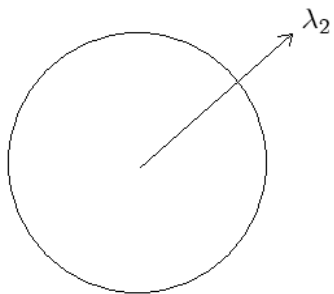
- 1  $E_t(M_{t+1}) = M_t$  for  $t < \bar{T}$  and  $t > \bar{T}$   
 $E_t(M_{t+1}) = 0$  for  $t = \bar{T}$

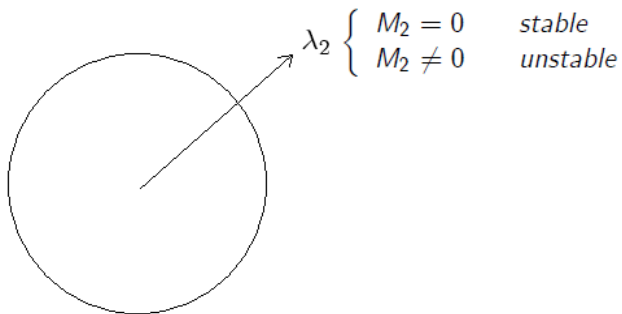
In general  $E_t(\eta_{t+1}) = 0$ .

- 2  $\lim_{h \rightarrow \infty} E_t M_{t+h} = 0$

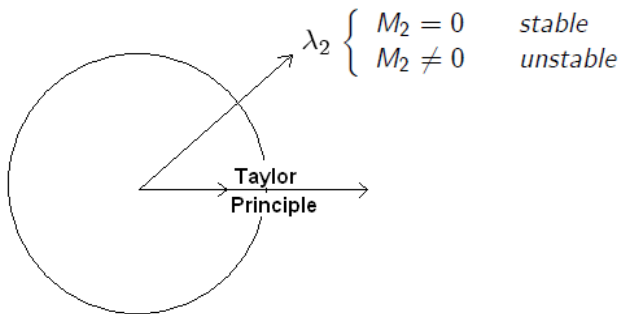
The transversality condition holds

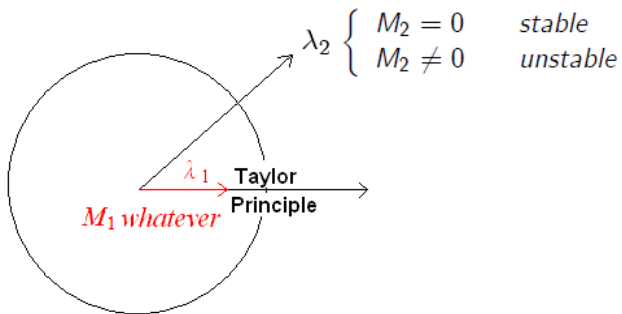
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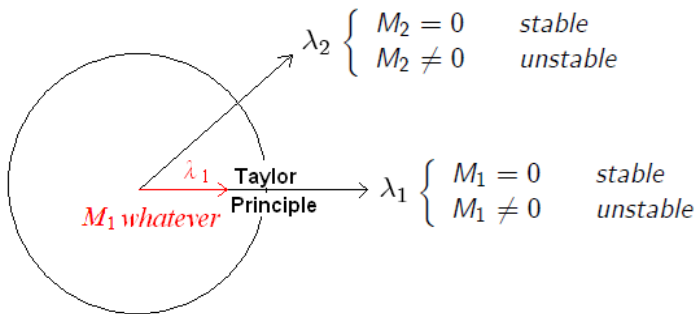


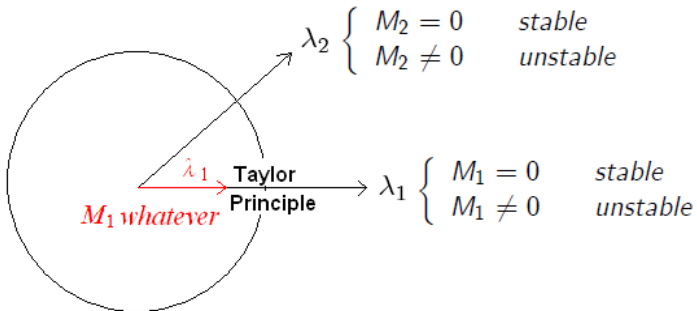












Compare two "models":  $M_S$  (stable solutions) and  $M_U$  (unstable solutions).

▶ Model

# Assumption on M

- UNDER STABILITY  $M_S$

If Taylor principle respected (determinacy)  $\Rightarrow M_t = 0 \forall t$

If Taylor principle **not** respected (indeterminacy):  $M_t = M_{t-1} + \zeta_t$

- UNDER INSTABILITY  $M_U$

$$M_t = N_t A_{t-1}$$

with

$$N_t = \begin{cases} N_{t-1}/\gamma + \zeta_t & \text{with probability } \gamma \\ 0 & \text{with probability } 1 - \gamma \end{cases}$$

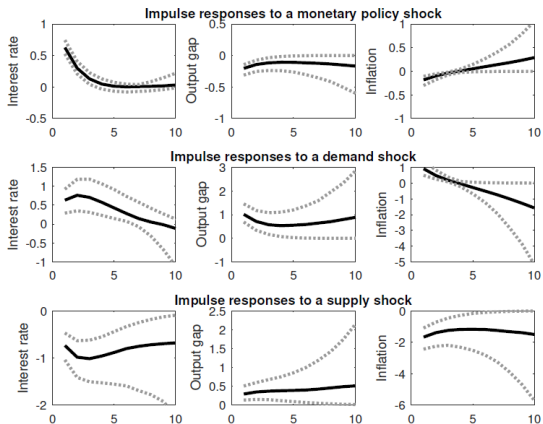
▶ Model

▶ MuGI

▶ MuGM

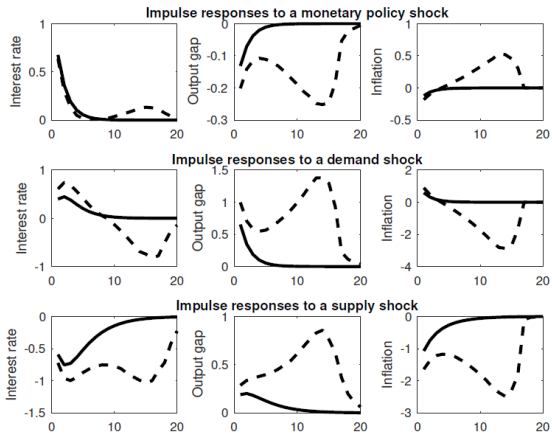
▶ priors

# Unstable Model, Great Inflation



**Transmission mechanism of structural shocks: GIRF in the  $M_U$  model**

# Unstable Model, Great Inflation



**Transmission mechanism of sunspot shock:** GIRF in the  $M_U$  model:  
solid line:  $M = 0$ , dashed line:  $M = 0.52$